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SKYLAB
SKYLAB S-019 FAR-UV
DATA (1979)
73-027A-02B

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SKYLAB S-019 FAR-UV

DATA (1979)

73-027A-02B

This data set catalog consists of one magnetic tape.
The tape is 9-track, 1600 BPI, ASCII with one file of data,
and was created on the 3081 computer. The D and C numbers
follow below.

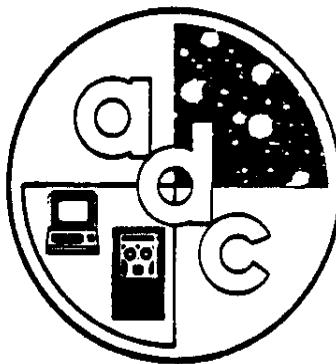
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ASTRONOMICAL DATA CENTER



DOCUMENTATION FOR THE
MACHINE-READABLE VERSION OF THE
CATALOGUE OF FAR-ULTRAVIOLET OBJECTIVE-PRISM
SPECTROPHOTOMETRY: SKYLAB EXPERIMENT S-019,
ULTRAVIOLET STELLAR ASTRONOMY

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DOCUMENTATION FOR THE MACHINE-READABLE VERSION OF
THE CATALOGUE OF FAR-ULTRAVIOLET
OBJECTIVE-PRISM SPECTROPHOTOMETRY:
SKYLAB EXPERIMENT S-019, ULTRAVIOLET
STELLAR ASTRONOMY

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SECTION 1 - INTRODUCTION

A COPY OF THIS DOCUMENT SHOULD ACCOMPANY EVERY DISTRIBUTED COPY OF THE MACHINE-READABLE CATALOGUE.

This document describes a machine-readable version of the Catalogue of Far-Ultraviolet Objective-Prism Spectrophotometry: Skylab Experiment S-019 (Henize et al. 1979). It contains arrays of ultraviolet flux measurements at wavelengths of 1300 to 4200 Å. The data were obtained with an objective-prism telescope (an f/3 Ritchey-Chretien system with a 15-cm aperture and a calcium fluoride-lithium fluoride focal-plane corrector). The spectra were digitized with a PDS 1010A microdensitometer, and each spectrum was scanned in a series of strips 30 microns wide apiece.

This catalogue contains data on 494 stars. The data on each star are quite complex and include a number of parameters besides the adopted fluxes. Near the beginning of Table 2-1 below is a description of the structure of the data on each star, and the user is urged to read this description carefully before trying to use the catalogue. The considerations that led to the adopted catalogue format are described in Section 4 below.

Those parts of the text of this document that are enclosed in quotation marks have been taken verbatim from Henize et al. (1979) for clarity.

REFERENCES:

Follow Section 4.

SECTION 2 - TAPE CONTENTS

A byte-by-byte description of the catalogue is given in Table 2-1 and Table 2-2. The information in the "Description" column is derived from the published catalogue as much as possible. The "Suggested Format" column is for Fortran formatted reads.

In Table 2-1, bytes can be numbered in two different ways: by position within a logical record, denoted by unsigned integers starting at 1; or by position within a field, denoted by signed integers, starting at +0. The first kind of number is called a byte number, and the second kind is called a byte offset, or just offset.

Offsets are used to describe field formats that are repeated many times in the record. Byte numbers are given (or can be computed) for the first byte of each instance of the field; then the byte numbers of the data within a particular field can be found by adding the offsets to that starting byte number.

Table 2-1. Tape Contents, Data File (1 of 9)
 Far-Ultraviolet Objective-Prism Spectrophotometry
 Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>	<u>SUGGESTED FORMAT</u>
1-10	Henry Draper Catalogue (HD) number of the star. Table 4-1 below is a list of all the stars in this file.	10A1
11-20	Alternative name of the star. When no alternative name is given, these bytes are the same as bytes 1-10.	10A1
	Each record in this file is 5354 bytes long and contains the data for one star. There are 494 stars, and thus, 494 records. The user must be aware of the general character of the record format before proceeding to the rest of Table 2-1.	
	The data in each record fall mainly into two categories: exposure data and flux data. Every record can contain data on as many as five different exposures, and data on as many as 514 different fluxes. Each exposure-data field is in one format, and each flux-data field is in another format. Thus, the format of each whole record consists, for the most part, of two series of repeated field types: one five-element series and one 514-element series.	
	The five exposure-data fields are contained in bytes 21-205, each field being 37 bytes long. The 514 flux-data fields are contained in bytes 214-5353, each field being 10 bytes long.	
	The series of flux-data fields falls into five sub-series, four of which correspond to constant wavelength intervals over which the fluxes are averaged. The fifth sub-series contains intermediate-band magnitudes averaged over intervals of varying width. In each of the five sub-series, a given wavelength interval is mapped into a certain location in the record, and this mapping is the same for the whole file. Not every sub-series is complete for every star, so that many flux fields are blank.	
21-205	Exposure data. These data are contained in five fields of 37 bytes apiece. For most stars, fewer than five exposures were used. In this case, the exposure-data fields are filled from left to right, and the fields not used contain blanks.	
	21 - 57 Data on first exposure. 58 - 94 Data on second exposure, or blank. 95 - 131 Data on third exposure, or blank. 132 - 168 Data on fourth exposure, or blank. 169 - 205 Data on fifth exposure, or blank.	

Table 2-1. Tape Contents, Data File (2 of 9)
Far-Ultraviolet Objective-Prism Spectrophotometry
Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>	<u>SUGGESTED FORMAT</u>
The exposure field format in terms of byte offsets (see page 2-1) is as follows:		
+0 to +6	Plate number. Information on each frame taken throughout the experiment can be found in the second file of this catalogue by keying on this plate number. See below Table 2-2, bytes 1-7.	7A1
+7	Blank.	1X
+8 to +10	Exposure time in seconds. If no precise time was available (because of gaps in the observing log), a reasonable default value is used.	13
+11	'.' if a default value is used for the exposure time in offsets +8 to +10, otherwise blank.	A1
+12	Blank.	1X
+13 to +23	Position of the star on the plate in millimeters (mm): "The plate position ... refers to the position of the optical head of the spectrum with respect to the field center, measured at the original plate scale. Coordinates are defined such that, with the frame oriented [with] shorter wavelengths toward the left, X increases toward the left and Y increases toward the bottom. The dimensions of the 4° by 5° field are 32 mm by 40 mm on this scale."	
NOTE: At this writing, prints of the S-019 frames can be obtained from the National Space Science Data Center as described in Section 4 below.		
+13 to +17	X (mm).	F5.1
+18	Blank.	1X
+19 to +23	Y (mm).	F5.1

Table 2-1. Tape Contents, Data File (3 of 9)
Far-Ultraviolet Objective-Prism Spectrophotometry
Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>	<u>SUGGESTED FORMAT</u>
+24	Blank.	1X
+25 to +26	Number of microdensitometer scans of the stellar spectrum. Each scan is a longitudinal strip 30 microns (μm) wide; thus multiplying by 30 μm gives the approximate width of the spectrum. The portions that overlap with other stars are not included.	I2
+27	Blank.	1X
+28 to +30	The weight of the spectrum. "The assigned weight ... for each spectrum depends partly on the measureable width and partly on the presumed quality of the reduction. The [Skylab mission] SL4 spectra normally receive a weight of 0.6 or 0.7, instead of 0.9 or 1.0, because the emulsion batch used was grainier and had poorer photometric properties than the batch used on SL2 and SL3."	F3.1
+31	Blank.	1X
+32 to +35	Scale value, which is "the relative scaling of the flux values for maximum agreement prior to averaging the different exposures. It also was used to make an approximate adjustment when less than the full width of the spectrum could be scanned."	F4.2
+36	Blank.	1X
206-213	Flux adjustment factor for the star.	
206	First part of comment notation. Can be "(", "<", or blank. See bytes 211-213.	A1
	The comment notation, which describes different cases and methods for the flux-adjustment calculation, is in bytes 206 and 211-213 in the same form as in the published catalogue. A full explanation of this notation is given in Section 4 below.	

Table 2-1. Tape Contents, Data File (4 of 9)
Far-Ultraviolet Objective-Prism Spectrophotometry
Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>	<u>SUGGESTED FORMAT</u>
207 - 210	Flux adjustment factor r, defined by this formula: (absolute flux) = (S-019 flux)*r	F4.2
	"The absolute flux levels for most stars were compared with other satellite measurements or with predicted fluxes to assess the accuracy of the calibration and to provide the user with adjustment factors."	
211 - 213	The second part of the comment notation, or else blank. The notations in this field consist of one or more of ")", ">", ":", "::", and "+-". See also byte 206.	3A1
214 - 5213	Flux data. Each flux is in a 10-byte field which contains all the information in the published catalogue except the wavelength. The wavelength is determined by the bytes in which the field appears. Note that floating-point (F format) data have implicit decimal points. The correspondence between wavelength and byte is determined by the following table, in which $\Delta\lambda$ is the bandwidth, MIN is the minimum of the possible wavelengths in angstroms (A) for that bandwidth, MAX is the maximum of the possible wavelengths in A for that bandwidth, and BASE is the first byte of the ten-byte field corresponding to MIN. The values MAX are redundant for determining the correspondence, but are given for clarity. COUNT is the greatest number of fluxes that can be given for each bandwidth; i.e., the number of ten-byte fields in that part of the record.	

Table 2-1(a). Constants for Determining Wavelength and Bandwidth from Byte Position.

<u>$\Delta\lambda$ (A)</u>	<u>BASE (BYTE)</u>	<u>MIN (A)</u>	<u>MAX (A)</u>	<u>COUNT</u>
2	214	1310	1828	260
5	2814	1800	2320	105
10	3864	2300	3040	75
20	4614	3000	4180	60

Table 2-1. Tape Contents, Data File (5 of 9)
 Far-Ultraviolet Objective-Prism Spectrophotometry
 Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>	<u>SUGGESTED FORMAT</u>
	<p>The wavelengths appear in ascending order starting at the given base bytes, each successive 10-byte field representing an increment by the bandwidth. Thus, the following formula converts a byte number (BYTNUM) to the corresponding wavelength (λ):</p> $\lambda = \lfloor (\text{BYTNUM}-\text{BASE})/10 \rfloor * \Delta\lambda + \text{MIN}$	
	<p>The following is a description of the flux field format in terms of byte offsets:</p>	
+0	<p>Comment character—either a blank or one of the following:</p> <ul style="list-style-type: none"> U - underexposure; the average intensity is only a little above the background fog E - extreme exposure; the average intensity is not far from saturation L - overlapping star D - plate defect 	A1
+1 to +4	<p>Flux measurement. Approximately the absolute flux incident at the earth, averaged over the range $\lambda - \Delta\lambda/2$ to $\lambda + \Delta\lambda/2$, where $\Delta\lambda$ has one of the values in Table 2-1(a), and where λ is related to the absolute position of the field in the record by the formula given after Table 2-1(a).</p> <p>This value is given in modified exponential notation, with two values D (digits) and C (order-of-magnitude code), such that</p> $\text{flux} = D * 10^{-C-8} \text{ ergs cm}^{-2} \text{ sec}^{-1} \text{ A}^{-1}$ <p>The flux is a weighted average of the data points for the wavelength from all the exposures. Each individual weight is a function of the slope of the characteristic curve of the emulsion at the density of the particular wavelength (which can vary from exposure to exposure), with the straight-line portion weighted at 1.0.</p>	

Table 2-1. Tape Contents, Data File (6 of 9)
Far-Ultraviolet Objective-Prism Spectrophotometry
Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>	<u>SUGGESTED FORMAT</u>
	There are eight dummy data points in the catalogue, and in the published version these are represented by flux values on the order of 10^{-24} . These instances are listed in Table 4-2 of this document and are represented by all-blank flux fields in the machine-readable version.	
+ 1 to + 3	Digits (D) of flux measurement, as in the above formula.	I3
+4	Order-of-magnitude code (C) of flux measurement, as in the above formula.	I1
+5 to +6	Total weight of the wavelength. This number is calculated from weights contributed by the individual spectra. It refers mainly to the accuracy of the absolute fluxes, rather than to spectral details. It also reflects the general reliability of the data in the region, as indicated by the comment character at offset +0. The weights of the component spectra are described under offsets +1 to +4.	F2.1
+7 to +9	Standard deviation, according to the following formula, as a percentage of the average flux given at offsets +1 to +4:	F3.1
	σ = the standard deviation N = the number of spectra averaged w_i = the weight of an individual spectrum as described under offsets +1 to +4 S_i = the scale factor for the exposure, given in offsets +22 to +25 of each exposure-data field (see bytes 21-205 above) F_i = the flux in an individual spectrum F = the average flux as given in offsets +1 to +4	

Table 2-1. Tape Contents, Data File (7 of 9)
Far-Ultraviolet Objective-Prism Spectrophotometry
Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>	<u>SUGGESTED FORMAT</u>
	$\sigma^2 = \sum_1^N w_i (S_i F_i - F)^2 / \sum_1^N w_i$	The standard deviations "reflect both grain noise and the effects of differences in slope among the derived energy distributions." A long series of zero standard deviations means that only one exposure was used in that region.
5214-5353	Intermediate-band magnitudes. These are computed over comparatively wide wavelength intervals and are related to the fluxes by	$MAG = -2.5 * \log F - 21.10$

where F is the average flux over the interval. The constant, corresponding to

$$3.64 \times 10^{-9} \text{ ergs cm}^{-2} \text{ sec}^{-1} \text{ A}^{-1}$$

was adopted in accordance with Nandy et al. (1976), "in order to put the magnitudes on the same energy scale as visual V magnitudes. The value of F is a straight average (rectangular passband), except for the 360 nanometer (nm) band, where a Johnson U filter function is used (Mathews and Sandage 1963). Based on the absolute calibration of U magnitudes (Johnson 1966), approximately 0.20 magnitude should be subtracted to put the ground-based U values on the same energy scale." In this magnitude system, the width of each passband on the film is a constant $360 \mu\text{m}$.

The intermediate-band magnitudes are in ten-byte fields of the same form as the flux fields above, with differences which will be explained below. The following table shows the starting byte of each intermediate-band magnitude field:

Table 2-1. Tape Contents, Data File (8 of 9)
 Far-Ultraviolet Objective-Prism Spectrophotometry
 Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>			<u>SUGGESTED FORMAT</u>
	FIRST BYTE	CENTRAL λ (nm)	RANGE (A)	FEATURES
Table 2-1(b). Intermediate-Width Passbands.				
5214	135	1347 - 1364		Continuum.
5224	139	1385 - 1405		SiIV lines.
5234	148	1465 - 1495		Continuum.
5244	154	1521 - 1560		C IV plus FeIII plus SiII lines.
5254	161	1587 - 1636		Temperature- and gravity-sensitive blends.
5264	166	1636 - 1693		Continuum; blends in hotter stars.
5274	172	1693 - 1760		Gravity-sensitive blend.
5284	181	1770 - 1853		Continuum.
5294	192	1870 - 1975		Gravity-sensitive depression (FeIII).
5304	204	1975 - 2110		Continuum.
5314	219	2110 - 2290		Continuum; maximum interstellar extinction.
5324	245	2320 - 2600		Continuum; FeII in cooler stars.
5334	280	2600 - 3070		Continuum; MgII in cooler stars.
5344	360	3070 - 4100		Continuum; approximate Johnson U magnitude

Following is a description of the field format of the intermediate-band magnitudes according to byte offsets.

+0	Comment character. See offset +0 under bytes 214-5213 above.	A1
+1 to +3	Magnitude as described above, with a decimal point implicit between offsets +1 and +2.	F3.2
+4	Not used.	1X
+5 to +6	Total weight for the passband. See offsets +5 to +6 under bytes 214-5213 above.	F2.1

Table 2-1. Tape Contents, Data File (9 of 9)
Far-Ultraviolet Objective-Prism Spectrophotometry
Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>	<u>SUGGESTED FORMAT</u>
+7 to +9	Standard deviation. See offsets +7 to +9 under bytes 214-5213 above. This standard deviation is to be interpreted slightly differently from those in bytes 214-5213: this value is "computed from the differences among the magnitudes from separate exposures, hence the grain noise component is eliminated."	F3.1
5354	Blank.	1X

Table 2-2. Tape Contents (1 of 4)
Exposure Data for Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>	<u>SUGGESTED FORMAT</u>
1-7	Frame number. Each exposure has a unique identifier which is referred to in the data file (see Table 2-1, bytes 21-205). Each exposure number, also called a frame number, consists of a mission number (SL2, SL3, or SL4), and an integer exposure sequence number.	7A1
8	Blank.	1X
9-21	Field center. "The celestial orientation of Skylab was often uncertain by 1° or more, so all field centers were determined with the Becvar atlases (Becvar 1962, 1964). On the prism exposures, the highly compressed optical end (head) of the spectrum, ending near 5000 Å, was used for positional reference." This field is blank if the frame was for calibration or was unusable because of defects, fogging, bad exposure, etc.	
9	"(" if the field center is uncertain, otherwise blank.	A1
10-14	Right ascension (1950) of field center.	
10-11	Hours of right ascension.	I2
12	A1	
13-14	Minutes of right ascension.	I2
15	Blank.	1X
16-20	Declination (1950) of field center in degrees.	F5.1
21	")" if the field center is uncertain, otherwise blank.	A1
22	Blank.	1X
23-27	Field designation. Five-character designation for internal use by the S-019 investigators.	5A1
28	Blank.	1X
29-34	Calendar date of exposure.	
29-30	Year.	I2
31-32	Month.	I2
33-34	Day.	I2

Table 2-2. Tape Contents (2 of 4)
Exposure Data for Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>	<u>SUGGESTED FORMAT</u>
35-36	Blank.	2X
37-48	Universal Time (UT) for the start of each exposure. Bytes 37-48 are blank for calibration exposures. "Start and end times could be determined to the nearest second from verbal 'marks' given by the astronaut and recorded on two-track tape, the second track containing a time signal. [Approximately 25 percent of] this information is lost, in which case the start time is given to the nearest minute and should be within 2 minutes of the actual start of the observation."	
	37-39 Day of year following January 0.0.	I3
40	":" (colon).	A1
41-42	UT of exposure, hours	I2
43	":" (colon).	A1
44-45	UT of exposure, minutes	I2
46	":", or blank if the exact start time was lost (see explanation of bytes 37-48).	A1
47-48	UT of exposure, seconds; or blank if the exact start time was lost (see explanation of bytes 37-48).	I2
49-50	Blank.	2X
51-53	Serial number of the film canister.	I3
54-55	Blank.	2X
56-59	Tilt angle, which refers to the Articulated Mirror System (Henize et al. 1979). At 0° tilt, the line of sight makes a 30° angle with the spacecraft wall. Bytes 56-77 are blank for calibration exposures.	F4.2
60-61	Blank.	2X
62-64	Position angle, defining the orientation of north on the frame. With the print oriented so that shorter wavelengths are toward the left, the zero point of position angle is toward the top. The angle is measured clockwise in degrees.	I3

Table 2-2. Tape Contents (3 of 4)
Exposure Data for Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>	<u>SUGGESTED FORMAT</u>	
65-66	Blank.	2X	
67-69	The planned, nominal exposure time in seconds.	I3	
70	Blank, or comment character: U = unwidened spectra N = no prism, direct photograph	A1	
71	Blank.	1X	
72	"(" if the actual exposure time is uncertain, otherwise blank.	A1	
73-75	Actual measured exposure time when determinable from the voice marks (see bytes 37-48).	I3	
76	")" if the actual exposure time is uncertain, otherwise blank.	A1	
77	Blank.	1X	
78-100	Remarks. Two exposures have extremely long remarks which are not included in the machine-readable version and are as follows:	23A1	
	<u>RECORD</u>		
	<u>NO.</u>	<u>FRAME</u>	
		<u>REMARK</u>	
	377	SL4-050	Started before comet rise, total 253 sec.
	417	SL4-090	Resembles very short U expo. According to transcript, it may have been terminated after 10 sec., <- 219.

Table 2-2. Tape Contents (4 of 4)
Exposure Data for Skylab Experiment S-019

<u>BYTES</u>	<u>DESCRIPTION</u>	<u>SUGGESTED FORMAT</u>
<u>REMARK</u>	<u>MEANING</u>	
ATM.EXT	Indicates spectra are affected by atmospheric extinction.	
END:	Gives end time of exposure to nearest second when known, if start time was not recorded.	
(FOGGED)	Indicates plate has narrow streak of fog, or general fog greater than normal.	
FOGGED	Indicates plate has heavy fog, but images may be usable.	
HIFOG	Indicates plate has severe fog, images not usable.	
NO IMD	Indicates the spectra are trailed due to failure to inhibit momentum dumping.	
<-	Indicates field was renumbered from the pad designation.	
->	Indicates this field now shared between two adjacent fields.	

SECTION 3 - TAPE CHARACTERISTICS

The information contained in this section describes the format of the catalogue as a whole, while Section 2 describes the format of the individual records. Multiple entries in the table refer to the respective files of a multiple-file catalogue. Besides the information given here, the user will need to know other attributes which depend on the specific operating system and hardware for which the catalogue tape has been generated. The names of these attributes are listed.

Table 3-1. Tape Characteristics

**Catalogue of Far-Ultraviolet Objective-Prism Spectrophotometry:
Objective-Prism Spectrophotometry: Skylab Experiment S-019,
Ultraviolet Stellar Astronomy**

Catalogue Abbreviation	S019
Number of Files	2
Logical Record Length (Bytes)	5354 and 100
Record Format (IBM JCL)	FB
Total Number of Logical Records	494 and 469

The other attributes of the tape will vary from copy to copy. They have been supplied to you separately, as they depend on the system for which the tape was made. The following space is provided for you to write them into this document.

Number of Tracks
Density (Bits per Inch)
Block Size (Bytes)
Blocking Factor (Number of Logical Records per Block)**
Total Number of Blocks
Character Code (BCD, ASCII, or EBCDIC)

**** Note:** The blocking factor is a fixed quantity for record format FB (fixed-block) files only. Even in FB files, the last block may be truncated.

SECTION 4 - REMARKS AND MODIFICATIONS

The S-019 catalogue was received from S. Parsons by the Astronomical Data Center (ADC), NASA Goddard Flight Center, on several tapes in 1979. The treatment of the exposure data file was straightforward, and will not be discussed here.

The spectrophotometric data was received in five files formatted for direct dumping to a printer to produce the catalogue as originally printed. The format has been radically changed. The data for each star have been put into a single record. In order to shorten the records to the degree possible, the flux data format has been condensed, mostly through the implicit representation of wavelength by the location of each datum in the record. No data have been changed or edited in regard to their content.

The format was designed in order to simplify the structure of any Fortran program that would use the data by making it possible for one Fortran READ statement to store in variables all of the required data for any given star. An example of Fortran code to accomplish this task is given in Section 6.

Prints of the S-019 fields can be obtained from the National Space Science Data Center. Address queries to the NSSDC, Manager, Request Coordination, Goddard Space Flight Center, Code 601, Greenbelt, MD 20771.

FLUX ADJUSTMENT FACTOR

This section describes the flux adjustment factor given in the first file (see Table 2-1, bytes 206-213). This explanation is partly quoted and partly paraphrased from the published catalogue.

"The [flux adjustment] factor is determined by comparison at selected wavelengths between S-019 fluxes resulting from the adopted calibration and one of the following:

- (a) The flux measured by the S2/68 spectrometer on the TD-1 satellite (Jamar et al. 1976; Willis and Wilson 1978).
- (b) The flux measured by the WEP spectrometer on OAO-2 (Code and Meade 1976), but adjusted by as much as 25 percent to agree with TD-1 fluxes on the average.

- (c) The flux computed from (i) intrinsic UV colors derived from TD-1 fluxes as a function of spectral type (Nandy et al. 1976), from (ii) the visual magnitude, and from (iii) interstellar extinction corrections using E(B-V) defined by the Q method from UBV photometry (Johnson 1958).
- (d) The flux from model atmospheres (Kurucz et al. 1974) and from visual magnitude."

Table 4-1 lists the notation in the machine-readable version of the catalogue for the above cases. See also Table 2-1, bytes 206-213.

Table 4-1. Flux Adjustment Factor Notation.

r is the flux adjustment factor given in bytes 207-210 of the data file (see Table 2-1). The associated notations, described in this table, are given in bytes 206 and 211-213. Cases (a) - (d) mentioned in this table refer to the description quoted in this section.

<u>NOTATION</u>	<u>DESCRIPTION</u>
r:	Adjustment factor determined for case (c). This method is restricted to spectral classes earlier than A0.
(r)	Adjustment factor determined for case (d). Used for stars in spectral classes A0 - A7 not measured by TD-1 or OAO-2.
r+-	Denotes a significant slope in the residual leading to a range on the order of 0.2 to 0.3 dex from the shortest to the longest wavelengths. A weighted mean r value is given.
<r>	The r value is predicted from appropriate shifts dependent on frame number and plate position for stars for which the r value could not be determined directly. See Henize <u>et al.</u> (1979), pp. 30-31, for more details.
r::	Adjustment factor determined for cases (c) and (d) when there is uncertainty about the appropriate intrinsic colors. The estimated color and predicted log r were averaged to obtain the final r value presented in the catalogue.

Table 4-2. Dummy Data Points.

This table refers to bytes 214 - 5213, offsets +1 to +4, in Table 2-1 above.

<u>FILE 1</u>			<u>NOMINAL FLUX</u>
<u>RECORD NO.</u>	<u>STAR</u>	<u>λ(A)</u>	<u>(ergs cm⁻² s⁻¹ A⁻¹)</u>
147	HD 55879	1384	1.23×10^{-24}
		1398	1.23×10^{-24}
		1402	1.23×10^{-24}
		1404	1.23×10^{-24}
171	HD 58978	1394	9.89×10^{-25}
		1404	9.89×10^{-25}
207	HD 69106	1550	1.50×10^{-24}
		1552	1.50×10^{-24}

Table 4-3. List of S-019 Stars. (1 of 4)

RECORD NUMBER	HD NUMBER	ALTERNATE NAME	RECORD NUMBER	HD NUMBER	ALTERNATE NAME
1	HD 358	ALF AND	71	HD 33948	HR 1704
2	HD 432	BET CAS	72	HD 34029	ALF AUS
3	HD 593	HD 593	73	HD 34035	BET ORI
4	HD 1337	ALO CAS	74	HD 34503	TAU ORI
5	HD 1976	HA 91	75	HD 34816	LAM LER
6	HD 2054	HA 90	76	HD 35039	HA 1764
7	HD 2772	LAM CAS	77	HD 35149	22 ORI
8	HD 2905	KAP CAS	78	HD 35299	23 ORI
9	HD 3240	HA 144	79	HD 35337	HA 1781
10	HD 3360	ZET CAS	80	HD 35439	8 LEP
11	HD 4142	HA 189	81	HD 35508	25 ORI
12	HD 4180	GMI CAS	82	HD 35575	GAM 3575
13	HD 5394	GAM CAS	83	HD 35608	HD 1803
14	HD 5406	HA 266	84	HD 35715	114 TAU
15	HD 10144	ALF ERI	85	HD 35912	PSI 1820
16	HD 10516	PHI PER	86	HD 36265	HA 1640
17	HD 16349	HD 16349	87	HD 36351	32 ORI
18	HD 18925	GAM PER	88	HD 36430	33 ORI
19	HD 19268	HR 930	89	HD 36436	IEZ UFS
20	HD 19350	BET PER	90	HD 36512	HA 1861
21	HD 20191	HD 20191	91	HD 36695	VV OH
22	HD 20263	HA 979	92	HD 36779	PHI-1 OH
23	HD 20677	32 PER	93	HD 36822	HD 36827
24	HD 20902	ALF PER	94	HD 36861	LAU OH
25	HD 21364	X4 TAU	95	HD 36935	HD 36895
26	HD 21428	34 PER	96	HD 37041	TOT ORI
27	HD 21551	HA 1051	97	HD 37128	EPS ORI
28	HD 21660	4 TAU	98	HD 37202	HK 1900
29	HD 21699	HA 1063	99	HD 37209	ZET TAJ
30	HD 21800	HA 1074	100	HD 37303	20 AUJ
31	HD 21933	6 TAU	101	HD 37401	21 SIG
32	HD 22192	ESI PER	102	HD 37490	22 CNG
33	HD 22780	HA 1113	103	HD 37507	23 ZET
34	HD 22928	DEA PER	104	HD 37742	24 ORI
35	HD 22951	40 PER	105	HD 37744	25 ORI
36	HD 23193	HA 1133	106	HD 37750	26 ORI
37	HD 23302	ETA TAU	107	HD 37756	27 ORI
38	HD 23630	HA 1172	108	HD 37765	28 ORI
39	HD 23753	42 PER	109	HD 37772	29 ORI
40	HD 23846	HR 1188	110	HD 37774	30 ORI
41	HD 23965	ZET PER	111	HD 37778	31 ORI
42	HD 24396	207 PER	112	HD 37787	32 ORI
43	HD 24504	X PER	113	HD 37793	33 ORI
44	HD 24534	Ha 1215	114	HD 37797	34 ORI
45	HD 24640	XI PER	115	HD 37798	35 ORI
46	HD 24912	40 PER	116	HD 37799	36 ORI
47	HD 25940	OMG TAU	117	HD 37807	37 ORI
48	HD 27045	53 TAU	118	HD 37813	38 ORI
49	HD 27295	50 TAU	119	HD 37817	39 ORI
50	HD 27309	53 PER	120	HD 37823	40 ORI
51	HD 27396	KAP TAU	121	HD 37827	41 ORI
52	HD 27934	67 TAU	122	HD 37833	42 ORI
53	HD 27946	UPS TAU	123	HD 37837	43 ORI
54	HD 28024	THT-2 TAU	124	HD 37844	44 ORI
55	HD 28319	61 TAU	125	HD 37848	45 ORI
56	HD 28548	83 TAU	126	HD 37855	46 ORI
57	HD 28879	HD 28879	127	HD 37860	47 ORI
58	HD 28910	AO TAU	128	HD 37866	48 ORI
59	HD 29479	SIG-1 TAU	129	HD 37872	49 ORI
60	HD 29488	SIG-2 TAU	130	HD 37877	50 ORI
61	HD 31512	62 ERI	131	HD 37883	51 ORI
62	HD 32060	ZET AUR	132	HD 37889	52 ORI
63	HD 32249	PSI ERI	133	HD 37895	53 ORI
64	HD 32630	ETA AUR	134	HD 37901	54 ORI
65	HD 32964	66 ERI	135	HD 37907	55 ORI
66	HD 33111	BET ERI	136	HD 37913	56 ORI
67	HD 33328	LAM ERI	137	HD 37919	57 ORI
68	HD 33904	MU LEP	138	HD 37925	58 ORI
69			139	HD 37931	59 ORI
70			140	HD 37938	60 ORI

Table 4-3. List of S-019 Stars. (2 of 4)

RECORD NUMBER	HD NUMBER	ALTERNATE NAME	RECORD NUMBER	HD NUMBER	ALTERNATE NAME
141	HD 53344	HD 53344	211	HD 69404	HD 69404
142	HD 54605	DEL CMA	212	HD 69973	HD 69973
143	HD 54893	HK 2702	213	HD 70309	HD 70309
144	HD 54912	HK 2704	214	HD 70556	HK 3283
145	HD 55522	26 CMA	215	HD 70930	HK 3294
146	HD 55657	HK 2734	216	HD 71935	HK 3350
147	HD 55879	HK 2739	217	HD 72108	HK 3363
148	HD 55908	HK 2741	218	HD 72232	HK 3386
149	HD 55905	HK 2743	219	HD 72737	HK 3385
150	HD 56014	27 CMA	220	HD 73105	HD 73105
151	HD 56139	OMI CMA	221	HD 74071	HK 34402
152	HD 56342	HK 2756	222	HD 74146	HK 3442
153	HD 56455	HK 2761	223	HD 74319	LD 74319
154	HD 56554	HD 56554	224	HD 74371	DEL VEL
155	HD 56779	ER 2770	225	HD 74956	HD 75241
156	HD 56970	HK 2774	226	HD 75241	HK 3527
157	HD 57000	TAU CMA	227	HD 75821	HD 76004
158	HD 57010	TAU CMA	228	HD 76004	HK 3562
159	HD 57150	HD 57193	229	HD 76506	HK 3582
160	HD 57219	HK 2790	230	HD 76728	HK 3593
161	HD 57393	HD 2800	231	HD 77002	HD 3600
162	HD 58011	HD 58011	232	HD 77320	HR 7804
163	HD 58260	HD 2823	233	HD 77475	A CAR
164	HD 58466	HD 2824	234	HD 78016	HK 361
165	HD 58325	ETA CMA	235	HD 79351	HK 3663
166	HD 58350	ETA CMA	236	HD 79410	HK 3672
167	HD 58420	ETA CMA	237	HD 79447	HK 3674
168	HD 58612	HK 2841	238	HD 79694	IOI CMA
169	HD 58978	HK 2845	239	HD 79735	HK 3850
170	HD 59020	HK 2855	240	HD 80404	HK 3863
171	HD 59138	HK 2860	241	HD 83979	HK 3868
172	HD 59247	HD 2870+71	242	HD 84226	HK 3883
173	HD 59550	HK 2873	243	HD 84609	HK 3892
174	HD 59604	HD 2884	244	HD 85871	VELOC
175	HD 59648	HK 2885	245	HD 86440	LEO LEO
176	HD 60098	HK 2895	246	HD 87737	ET ALF
177	HD 60312	HK 29344	247	HD 87901	41114 CAR
178	HD 60344	HK 29111	248	HD 90853	41180 CAR
179	HD 61071	HK 2937	249	HD 91310	41180 CAR
180	HD 61330	HK 2944	250	HD 91465	41205 CAR
181	HD 61429	HD 62315	251	HD 92664	4355 CEN
182	HD 62315	3 PUP	252	HD 92740	4492 CEN
183	HD 62623	HK 3004	253	HD 93030	MUS
184	HD 62747	HK 3025	254	HD 93194	LEO MUS
185	HD 63308	HD 63425	255	HD 93845	45490 CEN
186	HD 63425	OMI PUP	256	HD 97583	45490 CEN
187	HD 63465	HK 3035	257	HD 100841	4647 CEN
188	HD 63468	HD 63808	258	HD 101379	4647 CEN
189	HD 64305	HK 3078	259	HD 102249	4647 CEN
190	HD 64440	HK 3080	260	HD 102647	4647 CEN
191	HD 65315	HK 3107	261	HD 103079	4647 CEN
192	HD 66624	HK 3162	262	HD 106983	4647 CEN
193	HD 66811	ZET PUP	263	HD 108248	4823 CRU
194	HD 68092	HD 68092	264	HD 110335	4823 CRU
195	HD 68273	GAM VEL	265	HD 112078	4823 CRU
196	HD 68324	HK 3213	266	HD 113904	4975 MUS
197	HD 68357	HK 3227	267	HD 114529	4975 MUS
198	HD 68701	HD 68701	268	HD 114911	5093 VIB
199	HD 68895	HK 3234	269	HD 115846	5093 VIB
200	HD 69106	HK 3237	270	HD 116658	5206
201	HD 69144	HK 3244	271	HD 117651	5223
202	HD 69168	HK 3244	272	HD 120640	5223
203	HD 69302	HK 3250	273	HD 120991	52483 UPS-1 CEN
204			274		
205			275		
206			276		
207			277		
208			278		
209			279		
210			280		

Table 4-3. List of S-019 Stars. (3 of 4)

RECORDED NUMBER	HD NUMBER	ALTERNATE NAME	RECORDED NUMBER	HD NUMBER	ALTERNATE NAME
281	HD 124367	HA 5316	351	HD 150136	HA 6187
282	HD 125238	ICT LUP	352	HD 150166	HA 6188
283	HD 125286	HR 5358	353	HD 150898	HA 6219
284	HD 125721	HR 5375	354	HD 151515	HD 151515
285	HD 125823	HR 5378	355	HD 151804	HR 6245
286	HD 126341	TAU-1 LUP	356	HD 151890	MU-1 SCO
287	HD 126759	HD 126759	357	HD 151934	HR 6249
288	HD 126963	HR 5413	358	HD 152236	MU-2 SCO
289	HD 127331	SIG LUE	359	HD 152408	ZET-1 SCO
290	HD 127971	HR 5439	360	HD 152478	HR 6272
291	HD 127972	ETA CEN	361	HD 153261	HA 6274
292	HD 128345	KHO LUP	362	HD 153716	HA 6304
293	HD 128620	ALF CEN	363	HD 153919	HD 153919
294	HD 128974	HR 5466	364	HD 154090	HA 6320
295	HD 129052	ALF LUP	365	HD 155806	HA 6334
296	HD 129116	HD 129092	366	HD 155869	HD 155882
297	HD 129422	HA 5471	367	HD 156365	IOT ARA
298	HD 129929	HR 5482	368	HD 157042	THT OPH
299	HD 130559	HD 129929	369	HD 157056	44 OPH
300	HD 130701	MU LIB	370	HD 157792	HD 157832
301	HD 130807	AA CIR	371	HD 157832	HA 6490
302	HD 130819	OMI LUP	372	HD 157864	HA 6497
303	HD 130841	ALF-1 LIB	373	HD 157978	UPS SCO
304	HD 131120	ALF-2 LIB	374	HD 158408	51 OPH
305	HD 131492	HL 5543	375	HD 158643	HR 6622
306	HD 132058	THT CIR	376	HD 158704	LAM SCO
307	HD 132200	BET LUP	377	HD 158928	TH-1 SCO
308	HD 133738	KAP CEN	378	HD 159532	HR 6672
309	HD 133955	HD 133738	379	HD 162978	HA 6684
310	HD 134057	LAM LUP	380	HD 163472	HR 6716
311	HD 134067	HD 134657	381	HD 164402	68 OPH
312	HD 134067	HA 5651	382	HD 164447	93 SGK
313	HD 135160	HR 5661	383	HD 164577	96 HEA
314	HD 135240	DEL CIR	384	HD 164794	HD 165016
315	HD 135591	HA 5660	385	HD 164852	THT ARA
316	HD 135734	MU LUP	386	HD 165024	HD 165763
317	HD 135917	HD 135917	387	HD 165763	102 HEB
318	HD 136415	GAM CIR	388	HD 166182	MU SGK
319	HD 136504	EPS LUP	389	HD 166937	10 SGK
320	HD 138690	GAM LUP	390	HD 167263	10 SGK
321	HD 140008	PSI-2 LUP	391	HD 167264	HR 6875
322	HD 140784	HR 5860	392	HD 168905	EPS SGR
323	HD 141637	1 SCO	393	HD 169022	ALF TEL
324	HD 142114	2 SCO	394	HD 169407	DEL-1 TEL
325	HD 142165	HA 5906	395	HD 170405	HA 6960
326	HD 142184	HA 5907	396	HD 171034	ALF LYB
327	HD 142250	HA 5910	397	HD 172167	HR 7044
328	HD 142301	3 SCO	398	HD 173417	ZET LYB
329	HD 142883	HR 5934	399	HD 173646	LAM PAV
330	HD 142983	48 LIB	400	HD 173948	HR 7081
331	HD 142990	HR 5942	401	HD 174179	8 LYA
332	HD 143018	PI SCO	402	HD 174565	BET LYB
333	HD 143118	ETA LUP	403	HD 174638	ET-1 LYB
334	HD 143699	HR 5967	404	HD 174638	HA 7115
335	HD 144294	THT LUP	405	HD 174959	DEL-1 LYB
336	HD 144661	HR 5998	406	HD 175426	HD 175876
337	HD 145842	THT NOR	407	HD 175876	HR 7174
338	HD 147152	HR 6083	408	HD 176316	GAM LYB
339	HD 147165	SIG SCO	409	HD 176437	ZET AGL
340	HD 147894	HD 147894	410	HD 177724	BET-1 SGP
341	HD 147971	EPS NOR	411	HD 177724	BEI-2 SGP
342	HD 148478	ALF SCO	412	HD 181454	ALF SGR
343	HD 148605	22 SCO	413	HD 181623	BR 7392
344	HD 149038	MU NOR	414	HD 181809	V1264 CYG
345	HD 149404	HR 6164	415	HD 184905	14 CYG
346	HD 149438	TAU SCO	416	HD 185672	HD 186618
347	HD 149499	HR 6174	417	HD 186616	HR 7551
348	HD 149711	ZET OPH	418		
349	HD 149757	HD 150041	419		
350			420		

Table 4-3. List of S-019 Stars. (4 of 4)

RECORD NUMBER	HD NUMBER	ALTERNATE NAME	RECORD NUMBER	HD NUMBER	ALTERNATE NAME
421	HD 167879	V380 CYG	491	HD 222109	HR 8962
422	HD 188209	HB 7589	492	HD 222173	101 AND
423	HD 188252	HR 7591	493	HD 222439	KAP AND
424	HD 188439	V619 CYG	494	HD 224572	SIG CAS
425	HD 188892	22 CYG			
426	HD 189037	25 CYG			
427	HD 191010	28 CYG			
428	HD 192103	V1042 CYG			
429	HD 192105	HD 192163			
430	HD 192577	31 CYG			
431	HD 192909	32 CYG			
432	HD 193182	HD 193182			
433	HD 193237	P CYG			
434	HD 193365	36 CYG			
435	BD 193538	HB 7777			
436	HD 194335	HB 7807			
437	HD 197345	ALF CYG			
438	HD 199061	57 CYG			
439	HD 199579	HR 8023			
440	HD 200318	60 CYG			
441	HD 200595	HB 8064			
442	HD 201733	HB 8103			
443	HD 201819	HB 8105			
444	HD 202214	HB 8119			
445	HD 202347	HD 202347			
446	HD 202904	UPS CYG			
447	HD 203084	68 CYG			
448	HD 203260	ALF CEP			
449	HD 203330	HB 8164			
450	HD 203467	C CEP			
451	HD 204172	69 CYG			
452	HD 204403	70 CYG			
453	ED 205021	BET CEP			
454	HD 205139	HB 8243			
455	HD 205314	HB 8246			
456	HD 206165	9 CEP			
457	HD 206267	HB 8281			
458	HD 206365	HD 206365			
459	HD 206672	PI-1 CYG			
460	HD 206696	HD 206696			
461	HD 207330	PI-2 CYG			
462	HD 208682	HB 8375			
463	HD 208818	VV CEP			
464	HD 208947	HB 8384			
465	HD 209339	HB 8399			
466	HD 209461	14 CEP			
467	HD 209790	XI CEP			
468	HD 209975	19 CEP			
469	HD 210839	LAM CEP			
470	HD 211242	HB 8490			
471	HD 212124	2 LAC			
472	HD 212593	4 LAC			
473	HD 212883	HB 8549			
474	HD 212978	HB 8553			
475	HD 213310	5 LAC			
476	HD 214166	8 LAC			
477	HD 214263	HD 214263			
478	HD 214680	10 LAC			
479	HD 214993	12 DD LAC			
480	HD 216916	16 EN LAC			
481	HD 217050	EN LAC			
482	HD 217675	OMI AND			
483	HD 217943	HB 8777			
484	HD 218045	ALF PEG			
485	HD 218376	1 CAS			
486	HD 218440	HB 8803			
487	HD 218537	HB 8808			
488	HD 219634	HB 8854			
489	HD 220057	HD 220057			
490	HD 221253	AB CAS			

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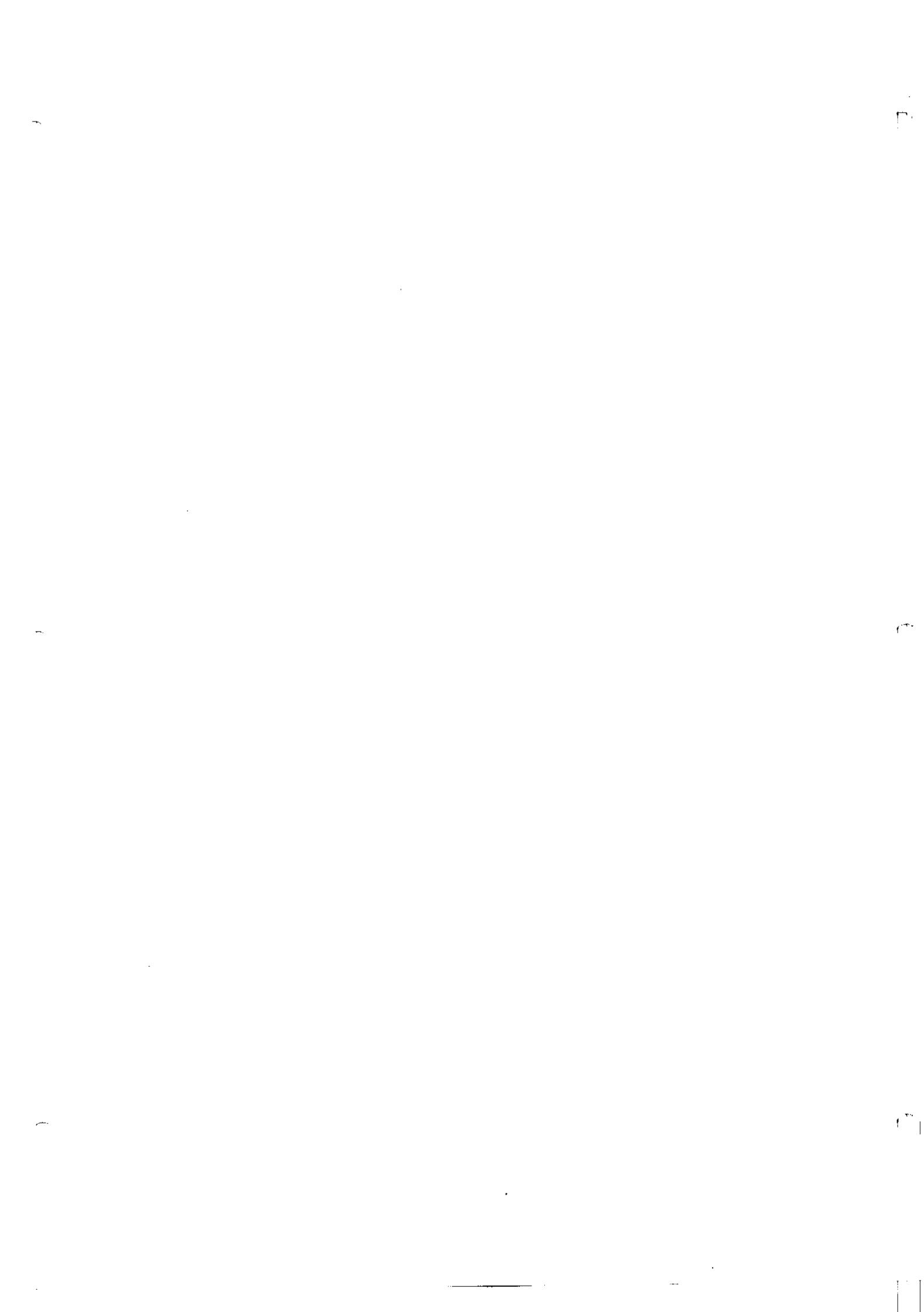
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SECTION 5 - SAMPLE LISTINGS

The following listings are intended for two purposes:

- (1) to be correlated with the record descriptions in Section 2 of this document;
- (2) to assist catalogue requestors in verifying that the request was filled properly.

The first lines of the listing underneath the heading are sequential numbers for the printed characters across the page. The numbers are to be read vertically. The bottom row contains the ones digits of the column numbers. For example:

11111111 ---
12345678901234567 ---
Column 1 Column 17

If each record has been split up into several lines for printing, this fact has been noted in the heading. Each line is printed starting at column 1 as defined by the column-number heading. (One byte as described in Section 2 above constitutes one character position in the listing.)

The heading lines and column numbers do not appear in the machine-readable file itself.

The listing for File 1 of the catalogue (described in Table 2-1) gives only the first record. The listing for File 2 (described in Table 2-2) gives the first one hundred records, and is made from a sample file having a blocking factor of 100 (see Section 3 above). The file you have received will very likely have another blocking factor.

First Record of Data File (1 of 2)

3019 CATALOGUE DATA FILE -- RECORD NO. 1

First Record of Data File (2 of 2)

First Block of Exposure File (1 of 5)

S019 EXPOSURE DATA FILE

LOGICAL RECORD LENGTH IS 100 BYTES
BLOCK SIZE IS 10000 BYTES
THERE ARE 100 RECORDS IN THIS BLOCK (BLOCKING FACTOR)
EACH RECORD IS PRINTED IN 2 LINES OF
65, AND 35 CHARACTERS

First Block of Exposure File (2 of 5)

First Block of Exposure File (3 of 5)

First Block of Exposure File (4 of 5)

First Block of Exposure File (5 of 5)

SECTION 6 - SAMPLE PROGRAM AND OUTPUT

A sample program to read in and print a portion of the S-019 catalogue is given in this section, along with the program's output for the first record in the catalogue. Those intending to use this program or portions of it should read the comment cards thoroughly first.

Sample Program (1 of 4)

```

PROGRAM S19PRT
SAMPLE PROGRAM TO PRINT OUT THE MACHINE-READABLE VERSION OF
THE S-019 SPECTROPHOTOMETRIC CATALOGUE DISTRIBUTED BY THE
ASTRONOMICAL DATA CENTER, NASA GODDARD SPACE FLIGHT CENTER.

THIS PROGRAM IS IN IBM FORTRAN 66. IT DOES NOT USE SPECIAL
DATA TYPES, SO IT SHOULD RUN ON MOST MACHINES WITH PEP. IF
ANY CHANGES ARE MADE TO THE PURPOSE OF THE PROGRAM, IT IS TO SHOW
HOW TO READ THE S-019 DATA FOR COMPUTATION. SOME DATA THAT
COULD HAVE BEEN READ IN AS CHARACTER STRINGS FOR
PRINTING ARE READ INTO COMPUTATIONAL VARIABLES INSTEAD.

FORTRAN UNIT 6 = PRINTED OUTPUT
FORTRAN UNIT 10 = S-019 CATALOGUE INPUT

WRITTEN BY R. S. HILL, SYSTEMS AND APPLIED SCIENCES CORP.-
21 NOVEMBER 1983

DECLARATIONS OF INPUT DATA. THE VARIABLES ARE GIVEN IN THE ORDER
IN WHICH THEY OCCUR IN THE RECORD. INTEGERS ARE USED TO HOLD
SINGLE CHARACTERS

STAR NAMES
  INTEGER NAME1(10), NAME2(10)
C..  DATA ON UP TO FIVE EXPOSURES
C..  INTEGER PLATE(7,5), EXPOS(5), EXPDEF(5)
      REAL ALEN(5), ANG(5)
      INTEGER SCANS(5)
      REAL WEIGHT(5), SCALE(5)

FLUX ADJUSTMENT FACTOR
  INTEGER COM1
  REAL K
  INTEGER COM2(3)

FLUXES AND MAGNITUDES
  INTEGER FCUM(500), D(500), C(500)
  REAL F, GT(500), SDEV(500)
  INTEGER MCUM(14)
  REAL MAG(14), MGT(14), MSDEV(14)

DECLARATIONS OF CONSTANTS FOR MAPPING THE SEQUENCE NUMBER OF

```

Sample Program (2 of 4)

```

C      A FLUX FIELD (1...514) TO A WAVELENGTH
C      INTEGER BNDWTH(4) /255 1020/
C      INTEGER MINWAV(4) /1316 1800 2300 3000/
C      INTEGER NFLUX(4) /260 10575 60/ 154, 161, 166, 172, 181, 192, 204, 219,
* 245, 280, 300/
C      THIS PROGRAM PROCESSES EACH RECOAD IN TURN AND TERMINATES
C      IF IT REACHES END-OF-FILE
C
C      KOUNT=0
100  READ(101,1) NAME1,NAME2,((PLATE(I,J),WEIGHT(I,J)),J=1,7),
*      EXPDEF(I),NAME(I),NAME(I),SCANS(I),
*      EXPDEF(I),NAME(I),NAME(I),SCANS(I),
*      EXPDEF(I),NAME(I),NAME(I),SCANS(I),
*      EXPDEF(I),NAME(I),NAME(I),SCANS(I),
*      EXPDEF(I),NAME(I),NAME(I),SCANS(I)
1    FORMAT(10A1,10A1,5(7A1,1X),13A1,1X,F4.2,A1,225(1A1,F2.1),
*      14(A1,F3.2),14(A1,F2.1,F3.1)).
C
C      THIS PROGRAM ENDS AFTER PROCESSING FIVE STARS; THE WHOLE
C      CATALOGUE WOULD PRODUCE MANY MANY LINES OF OUTPUT. THE
C      PROGRAM CAN BE MODIFIED EASILY TO PRINT A PARTICULAR RECORD
C      OR SEQUENCE OF RECORDS; SEE TABLE 4-3 FOR SECTION PRODUCED BY THIS PROGRAM FOR THE FIRST STAR ONLY, ALPHABETICAL ORDER.
C
C      KOUNT=KOUNT+1
IF (KOUNT .GT. 5) GO TO 999
C      PRINT STAR NAME, EXPOSURE DATA, AND FLUX ADJUSTMENT FACTOR
C
C      WRITE(6,2),NAME1,NAME2,* *** BEGINNING OF DATA FOR STAR *** .
2    * 3A10A1
      WRITE(6,3)
3    FORMAT(6,3) PLATE(I,J) EXP TIME X(MM) Y(MM) SCANS(I),
*      WEIGHT(SCALE/I)
DO 150 I=1,5
IF (EXPDEF(I)=0) EXIT: AND, WEIGHT(I,J)=0) GO TO 150
      WRITE(6,4) (PLATE(I,J),WEIGHT(I,J),EXPDEF(I),NAME(I),SCANS(I),
*      EXPDEF(I),NAME(I),NAME(I),SCANS(I),
*      EXPDEF(I),NAME(I),NAME(I),SCANS(I),
*      EXPDEF(I),NAME(I),NAME(I),SCANS(I),
*      EXPDEF(I),NAME(I),NAME(I),SCANS(I))
150 CONTINUE
      WRITE(6,5) COM1,RCOM2
      5 FORMAT(6,5) FLUX ADJUSTMENT FACTOR = ' ,A1,F4.2,3A1/)

C      PRINT FLUXES 1 DEG ON A LINE-OUT TABLE AND TWO MAIN APPROXES
C      TO THIS PROBLEM. ONE IS TO BUILD AN OUTPUT LINE AND PRINT AT
C      THE OTHER IS TO USE THE OVERSTYLIC CATALOGUE CHARACTER:-

```

Sample Program (3 of 4)

```

C THIS PROGRAM USES OVERSTRIKES
C
C   WRITE (6,6) LAMBDA  FLUX  WEIGHT ST-DEV  "
C   *          LAMBDA  FLUX  WEIGHT ST-DEV  "
C   *          LAMBDA  FLUX  WEIGHT ST-DEV  "
C
C..  LOOP THROUGH THE FOUR GROUPS OF NAKOM-BAND FLUXES
C
C..  LDONE=0
DO 500 I=1,4
LDONE=LDONE+1
LST=LDONE+NFLUXA(I)
INTBND=BND*TH(I)
LAMBDA=MINTAV(I)
LINE=1
C
C..  LOOP THROUGH THE FLUXES ASSOCIATED WITH EACH BANDWIDTH
C..  BLANK FIELDS (CONVERTED WHEN READ TO ZEROES). INDICATE NO
C..  DATA AT THAT WAVELENGTH
DO 300 J=LST,LST
IF (D(J)/100.0 .EQ. 0) GO TO 300
FLX=D(J)/100.0
IEXP=-6-C(J)
GO TO 210
210 WRITE (6,15) LAMBDA,FCON(J),FLX,IEXP,FNGT(J),SDEV(J)
GO TO 210
210 FORMAT (1A,14,A1,3X,F4.2,E,13,3X,F4.1,5X)
GO TO 250
250 WRITE (6,20) LAMBDA,FCON(J),FLX,IEXP,FNGT(J),SDEV(J)
GO TO 250
250 FORMAT (1A,14,A1,3X,F4.2,E,13,3X,F4.1,5X)
GO TO 250
250 WRITE (6,30) LAMBDA,FCON(J),FLX,IEXP,FNGT(J),SDEV(J)
GO TO 250
250 FORMAT (1A,14,A1,3X,F4.2,E,13,3X,F4.1,5X)
LINE=LINE+1
IF (LINE>LST) LINE=1
IF (LAMBDA=LAMBDA+INTBND)
      WRITE (6,50)
      50 FORMAT (1A,14,A1,3X,F4.2,E,13,3X,F4.1,5X)
      500 FORMATT(1A,14,A1,3X,F4.2,E,13,3X,F4.1,5X)
      LDONE=LST
C
C..  PRINT INTERMEDIATE-BAND MAGNITUDES
C
C..  WRITE (6,6) LAMBDA  MAGNITUDE  WEIGHT ST-DEV "
C..  60 FORMAT (1A,14)
DO 600 I=1,14
IF (MAG(I) .EQ. 0.0 .AND. MGT(I) .EQ. 0.0 .AND. MSDEV(I)
* .EQ. 0.0) GO TO 600
* WRITE (6,6) INTBND(I),MCOM(I)
* MAG(I),MGT(I),MSDEV(I)
600 FORMATT(1A,14)

```

Sample Program (4 of 4)

```
61 FORMAT(3A,13,A1,6A,F4.2,F4.1,6X,F4.1)
600 CONTINUE
C...    END OF STAR
C      WRITE(6,92) NAME1,NAME2
C      FORMAT(//1X,10A1,3X,*** END OF DATA FOR STAR *** ,3X,10A1)
C      GO TO 100
C...
C...    EXIT FROM PROGRAM
C 999 STOP
END
```

Output of Sample Program (1 of 3)

Output of Sample Program (2 of 3)

Output of Sample Program (3 of 3)

בְּאֶלְעָזָר בֶּן־בְּנֵי־יִשְׂרָאֵל

Digitized by srujanika@gmail.com

କାନ୍ତିରେ ପାଦ ପାଦରେ କାନ୍ତି
କାନ୍ତିରେ ପାଦ ପାଦରେ କାନ୍ତି

בְּרֵבָדָה וְבְרֵבָדָה

ପାତ୍ରମାନଙ୍କ ପାତ୍ରମାନଙ୍କ

LATITUDE	HEIGHT	ST-DIST
13.9	0.71	1.6
13.9	0.63	1.4
13.9	0.55	1.2
13.9	0.47	1.0
13.9	0.39	0.8
13.9	0.31	0.6
13.9	0.23	0.4
13.9	0.15	0.2
13.9	0.07	0.1
13.9	-0.01	0.05

358 *SEE THE DATA FOR STAIRS* ALP AND QR

***** LIST OF 1ST & LAST RECORD ON SO-1 *****

INPUT PARAMETERS ARE: AS FL=1±1 1

TAPE NO.	1	FILE NO.	1
RECORD	1	LENGTH	26770
HDO	358 ALF AND	SL4-	67 217
69	28 - 10.5	.7	-10.5
		19	.8 .97
			1.2345

	U1523	4	63U2023	9	45U2273	5	3U2
133	5209U1973	6	84U2153	6	6 2163 6 43U1963	6194U1773	7 12 2343 9 47 2453 9344 2243 9370 2183 9207
1513	9236	211310244	221310200	193310203	190310 49	176310144	186310234 174310199 1943 9 18 1613 9 83
1463	9	82 1443 9 12	1343 8249	1363 9155	172310119	205311106	209311 66 226311 12 226311 1
3	227313	53 2123012 38	206312 79	195312 27	213313 27	213314 58	225314 105 224314 59 221314 73 208314 59 221315
13	218313	77 220313	27	213314 56	241314 56	225314 105 224314 59 221314 73 208314 59 221315	
82	206315	82 198315104	194315104	211314133	205314129	208314121	201315128 185315102 185316146 20231
6106	200316130	219315178	207315172	208314192	209315206	201315188	194316191 203318153 206319102 2003
19	19 68	207319100	1953100	195318 99	189317147	194317168	197318118 195319 83 197319 77 188318196 1803
318151	176317164	156317164	158317176	166317184	177318180	172318155	178318 84 203319 45 196319 71 199319
4318	32 200318	23 198318130	195318145	204316 88	208317 98	203317135	192317140 198317 88 206318126 2
8318112	201318 98	204318111	202318136	200318165	201319 97	202319 58	207320 10 211320 27 214319 27
219319	93 227319	89 227319	83 218319	90 223319	86 230319	71 227319	66 222319106 213319114 205319 76
212319	49 205319	42 196319	64 196319	54 205319	62 166319	54 214319 78 213319 51 220319 3	213319 51 220319 3
66	226319	81 218319	81 210319	19 210319	19 210319	46 206319 53 202319	36 198319 36 198319 71 199319
45	210319 91	61 217319	80 217319	89 209319	92 209319	86 200319 86 200319	79 204319 49 204319 20 208319 47 219319
100	222319131	216319130	214318 86	221318 88	229318131	229318134	228318133 236318120 234318 72 22331
66	216319	62 215319	33 213319	33 206319	69 197319	81 196319	23 204318 30 212318 8 207318 19 1993
18	10 196318	34 190319	33 185319	33 185319	56 185319	83 183319 48 180318 52 180318 72 179318	64 179318 81 173318 16 179318 16 179318 60 1
18	19 86	16631719	66 167319	51 173319	45 173318	52 174318 65 174318 56 174318 27 169318	16 169318 16 169318 60 1
6318	64 182317 58	186317 86	185317115	181317120	179317102	176317 84 168317	83 160317 83 160317 102 156318115 1
6718	92 157318	80 156318	65 151318	62 143318	65 139318	78 139318 94	142318 82 146318 82 148318 99
145318108	141318106	141318105	144317 98	149317 88	152317 76	152317 75 150317	79 150317 83 152317 87
155317	98 157317	97 157316	98 156316110	154316109	150317106	145317104	142317107 143317 94 147316 7
150316	69 146316 66	147316 71	145316 66	152317 84	156317 95	153317101	104 143316102 147316 71 147316 71 147316 71
70	142316 80	149316 81	143316112	127316102	125316 81	131316113	135316109 143316109 143316102 156318115 1
1118	128316 94	126316 38	126315 68	126315 98	135315113	134315 79	126315 70 128315154 128315123 13231
5134	128315154	124315113	122314 88	123314 88	123314 52	127314 19	128314 20 129314 54 132313105 135313102 13231
13	73 132313112	140313 87	141313112	139312 77	137312 63	134312 54 126312 38	122312 72 124312 29 129
5132	45 123311 30	123311 41	123311 41	123311 63	129311 66	127311 51 131311 71	133311 51 131311 71 131311 71 131311 71
7311	31 110311 25	110311 58	118311166	119310204	116310204	109310168	104310131 110310131 111310146 1
143	9131 1163	9122 1203	8167 1233	8208 1283	8181 1293	8123 1223	8127 1143 8174 1113 8188 1083 8183
1033	8209 9974	8222 1003	8209 1023	8149 1043 8	90 1093 7	97E1123 7154E1143 7192E1143 7207E1133	7228
1103	7244E1083	7236E1053	7204E1023	7179 9764	7166 9404	7162 9084	7183 8904 7241 8954 7304E9214 735
35E9474	735E95904	7419E8744	7411E8734	7379E87204	7352 7864	8329 7704 8285 7624	8285 7624 8285 7624 8285 7624 8285 7624
974	8308 7644	8243 7604	7424E8124	7270E8434	7281E8374	6332E8164	6417E804 6427E7574 7343E6984 7
3116E544	7312 6244	7277E6554	7274E6854	7268E6684	7251E6074	7260 5604	7282 5614 7301E5664 7329E5734
77338E5814	7328E5794	6333E5904	6324E6154	6292E6294	6266E6334	6259E6384	5279E6284 5308E5984 5315E5674
6317E5504	6303E5494	6293E5524	5305E5524	5314E5434	5287E5334	5240E5304	5232E5224 5264E5134 5270E5094
4	5229E5494	5198E5104	5221E4734	6229E4734	6289E4734	6277E4734	6285E4734 6285E4734 6285E4734 6285E4734 6285E4734
74	5371E4584	5381E4424	5392E4204	5381E4074	5344E4034	5309E4034	5281E4034 5224E4034 5224E4034 5224E4034 5224E4034
104	5139E4054	5133E3934	5139E3784	5174E3604	5226E3464	5275E3424	5297E3504 5293E3664 4259E3584 4211E3524 4256
4044	4150E4114	494E4084	461E3964	476E3814	4125	E4114 4 95E3964 4 83E3684 4184E3524 4256	
3434	4291E3444	4286E3464	4240E3414	4200E3384	4194E3384	4156E3264	4191E3064 4174E3024 4195E3134 419
5532	5532E2504	4 93E2524	4 92E2524	4 92E2524	4 92E2524	4 92E2524	4 92E2524 4 92E2524 4 92E2524 4 92E2524
1817	83174	HD 4312	BFI 57	878 2255	901 S1	91 S1	91 S1 712 5 11 5281E5204 5281E5204 5281E5204 5281E5204 5281E5204

1 0004575 2 0003855 1 0003575 1 0004285 2 0005645 3 0006205 3 0005435 3 0005155 3 0005575 3 0005725
 4 0005555 4 0005305 4 0005615 4 00 5995 4 0005915 4 0005575 4 0005235 4 0004805 4 0004705 4 00 549
 5 5 00 5985 6 00 5075 5 0004285 4 0004215 3 0004145 3 0004225 3 0005075 4 00 6065 6 00 6095 6 00 57
 35 6 00 5875 6 00 6325 7 00 6605 8 00 6225 7 00 5905 7 00 6155 8 00 681510 00 731510 00 744510 00 7
 54510 00 753510 00 737510 00 729510 00 751510 00 786510 00 793510 00 791510 00 799510 00 805510 00
 786510 00 766510 00 792510 00 857510 00 752510 00 735510 00 779510 00 796510 00 786510 00
 826510 00 907510 00 892510 11 877511 50 817511 85 761511 74 678511 11 716511 51 712512 70 726512 3
 3 759512 26 800513 77 827513 58 819513 17 851514 45 904515 3 942516 27 970517 4 957517 54 932517
 26 958518 46 943518 64 928519 30 934519 11 100420 58 106420 54 111420 25 1117420 33 120420 25 121419
 19 123419 34 122419 7 118419 17 117419 5 115419 19 106419 46 998519 70 102419 50 103419 54 10341
 9 52 106418 8 108418 9 111418 5 116418 24 125417 29 131417 38 132416 25 134416 10 137415 14 1444
 15 23 151414 26 157414 2 162413 12 164413 2 161413 9 157413 22 149413 20 145413 9 144413 10 142
 413 23 140413 19 140412 17 141412 12 139412 17 136412 17 134412 8 132412 2 129412 1 126412 14 12
 5412 42 126411 74 131411 64 138411 40 145410 23 151410 25 153410 44 1514 9 38E1514 9 19E1504 9 19E1
 494 9 31E1504 9 36E1524 9 16E1534 9 4E1494 9 10E1464 9 16E1454 9 3E1464 8 16E1454 8 28E1434 8 26E
 1404 8 28E1394 8 40E1404 8 44E1424 8 39E1454 8 32E1454 8 22E1454 8 22E1424 8 27E1394 8 17E1364 8 5
 E1454 8 22E1394 8 16E1314 8 9E1204 8 31E1114 8 95E1064 9 105E9935 9 70E9445 9 47E9245 9 6
 6E9205 9 45E9535 8 43E9725 8 61E1024 8 50E1104 7 40E1144 7 28E11204 6 44E1184 6 23E1204 6
 10E1234 5 3E1244 5 14E1174 5 27E1164 5 17E1184 5 9E1214 5 3E1264 4 3E1334 4 6E1364 4
 12E1414 3 5E1424 3 4E1484 3 18E1534 3 12E1604 3 21E1664 2 6E1724 2 12E1884 2 10E1954 2 12E1994
 2 3E1904 2 3

0 00 000 00 00 000 00 00 000 00 000 00 000 00 000 00 000 00 000 00 000 00 000 00 000 00 000 00 000 00 000 00
 0 00 000 00 00 00 HD 593 HD 593 SL3- 78 225: 9.8 -17.8 22.1.0 .97 SL3- 79 77: 9.8 -17.8 2
 1.0 1.04 0.98

33 326417155 294417134 280417173 309417172 316417160 318417121 330417 94 353417 87 320417 79 303417
 7 327417 22 375417 49 386417 37 399417 19 435417 19 397417 22 329417 64 327417114 329417108 32441
 7 78 317417 82 345417 61 351417 41 350417 47 383417 47 393417 79 372417 82 359417 25 371417 77 3604
 17 39 362417 47 369417 55 347417 81 341417 65 357417 71 367417 42 381417 24 390417 73 405417 67 415
 417 41 401417 42 389417 16 377417 57 374417 16 405416 19 426416 32 433416 24 438416 11 437416 10 44
 3416 45 444416 34 438416 3 423416 14 417416 27 421416 16 431416 49 434416 29 437416 30 439416 81 4
 45416 96 467416 52 492416 22 492416 52 464416 5 446416 20 471416 12 496416 67 488416 76 477416 35
 481416 15 487416 25 478416 9 465416 32 475416 54 490416 65 472416 54 454416 22 442416 16 454416 18
 459416 18 452416 31 440416 55 441415 50 450415 43 457415 67 462415 87 476415 78 490415 63 489415 3
 4 484415 12 475415 8 461415 4 443415 35 435415 65 430415 79 423415 45 434415 12 454415 14 468415
 29 485415 29 515415 8 528415 22 516415 21 494415 20 474415 57 461415 72 473415 56 489415 45 492415
 50 50 476415 52 462414 60 464414 75 471414 85 473414 86 461414 84 446414 76 455414 97 466414 102 48041
 4110 489414 88 484414 52 475414 31 462414 15 457414 28 458414 26 471414 23 480414 36 4894
 14 72 494414 85 497414 69 483414 59 469414 75 447414 65 481414 61 457414 19 458414 20 483
 414 53 498414 82 458414 76 438414 11 454414 34 454414 14 429414 79 435414 47 429414 25 447414 26 45
 8414 46 419414 17 419414 25 413414 27 379414 38 361414 58 363414 28 371414 38 372414 34 366414 21 323414 3
 582413 14 35614 60 361414 3 364414 34 375414 84 352414 61 319414 59 315414 21 319414 21 323414 46
 32413 27 324413 25 298413 17 305413 17 328413 45 330413 87 312413 62 301413 49 298413 46 293412 16
 293412 41 304412 26 310412 33 287412 6 274412 65 274412 48 273412 21 266412 18 275411 30 295411 5
 4 287411 11 283411 43 274411 64 260411 13 275411 18 277411 6 262410 30 254410 29 259410 18 270410
 8 275410 2 273410 8 283410 24 2824 9 51 2664 9 41 2564 9 31 2564 9 33 2594 9 45 2534 9 30 2554 9
 10 2554 9 0 2514 9 13 2554 9 2 2554 9 2 2474 9 4 2464 9 41 2554 9 101E2594 8109E2694 8102E2794
 8 76E2694 8 14E2564 8 26E2504 8 22E2504 8 9E2524 8 38E2514 8 42E2584 8 54E2614 7 42E2694 7 46E2744
 7 32E2744 7 5E2684 7 27E2774 7 3E2914 7 21E2934 7 22E2944 7 32E2824 6 31E2784 6 65E2754 6 88E276
 4 6 90E2764 6 75 E2524 6 8E2754 6 79E2744 6 61E2794 6 47E2844 6 23E2834 6 52E2754 6 80E26
 94 6 83E2734 5 96E2874 5 5122E2914 5 75E2994 5 38E3034 4 66E3024 4 87E3054 4 78E3054 4 67E3174 4 83E3
 144 4 90E2954 4 53E2834 3 18E2954 3 11E3244 3 31E3364 3 61E3234 3 78E3104 3 86E3144 3 94E3244 3 90E
 3184 3 59E3034 3 23E2954 2 7E2974 2 00E3054 2 00E3194 2 00E3244 2 00E3194 2 00E3134 2 00
 E3064 2 00E3034 2 00E3094 2 00E3274 2 00E3534 2 00E3414 2 00E3244 2 00E3214 2 00E3204 2 00
 DE3304 2 00E3524 2 00E3774 2 00E3764 2 00E3564 2 00E3284 2 00E3294 2 00E3404 2 00E3394 2

U249 2 00 250 12154 273 16174 250 17 14 230 16 19 223 15 29 223 14 56 252 14 26 276 12 70 284 8 8
 9E271 4 71 000 00 00

**** JOB DONE.
 SWO LPS

SNOP *****	TDISCI *****	*****	TDISCI *****	*****	TDISCI *****
SNOP *****	TDISCI *****	*****	TDISCI *****	*****	TDISCI *****